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High-Frequency Plasma Conductivity

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ABSTRACT

On the basis of the BBGKY hierarchy of equations an expression is derived for the response of a fully ionized plasma to a strong, high-frequency electric field in the limit of infinite ion mass. It is found that even in this limit the ion-ion correlation function is substantially affected by the field. The corrections to earlier nonlinear results for the current density appear to be quite essential. The validity of the model introduced by Dawson and Oberman to study the response to a vanishingly small field is confirmed for larger values of the field when the correct expression for the ion-ion correlations is introduced; the model by itself does not yield such an expression. The results have interest for the heating of the plasma and for the propagation of a strong electromagnetic wave through the plasma. The theory seems to be valid for any field intensity for which the plasma is stable.

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I. INTRODUCTION

The high-frequency conductivity of a fully ionized plasma was first properly computed by Dawson and Oberman,^{1,2} who used a simple model for a plasma of infinitely heavy ions. In this limit, they regarded the ions as a set of fixed scatterers; first the ions were assumed to be randomly distributed¹ and later ion-ion thermal equilibrium correlations were taken into account.² Although only the linear limit (when the driving field is taken to be vanishingly small) was considered, it was realized that the model in principle could be used to study nonlinear effects. It was essential to bring into the model, as an external datum, an expression for the ion-ion correlation function.

In a parallel development it was shown³ that a linearized treatment of the BBGKY hierarchy gives the same result obtained in Ref. 2; the thermal ion-ion correlations were introduced from the outset. It may be concluded that, in the linear limit, the Dawson-Oberman model gives correct results. There is actually no problem about choosing a proper expression for the ion-ion correlations because the plasma is near thermal equilibrium; moreover, no substantial difference was found even between the results of Refs. 1 and 2.

In the present paper we use the BBGKY hierarchy of equations to obtain the plasma response to a stronger, high-frequency electric field. While the concept of conductivity, strictly, is defined only for vanishingly small fields, an expression for the current is needed nonetheless to study the nonlinear absorption of the field energy by the plasma and the propagation of a ~~strong~~ electromagnetic wave.

Nonlinear effects have been considered by Silin,⁴ who used the BBGKY equations but failed to include collective phenomena, and Albin and Rand,⁵ who assumed a driving frequency, ω , large compared with the electron plasma frequency, ω_p . In both these papers the typical resonant plasma effects¹ were missed. Kaw and Salat⁶ extended the Dawson-Oberman model to moderate fields (in the sense that the resulting directed component of electron velocity is smaller than the electron thermal velocity) and found new resonances in the current. In both Refs. 5 and 6 the assumption of uncorrelated ions was made. While the model has been shown to be valid only in the linear limit, one would feel its validity could be extended to larger fields (a first result of our study is, indeed, the confirmation of this validity). The real difficulty in using the model for nonlinear considerations lies in the selection of a proper ion-ion correlation function.

The limit of an infinite ion mass, m_i , was taken in Ref. 6, as will be the case here. It could be thought that in this limit the field would not modify the ion-ion correlations, because there would be no response to the field from the infinitely heavy ions, thereby allowing the use of the thermal correlations, i. e., the correlations existing before the field had been switched on. This consideration (together with the facts that no substantial difference was found between the results of Refs. 1 and 2 and that computations would be somewhat simpler for uncorrelated ions than for thermal correlations) would seem to justify the selection of zero ion-ion correlations made in Ref. 6. The most important results of the present investigation are

(1) the conclusion that the argument just given is wrong and (2) the derivation of an expression for the ion-ion correlations quite different from the thermal value (but to which it reduces for very weak fields). The expression contains additional resonances, thus modifying substantially in some cases the results of Ref. 6 .

The change in the correlations, in the limit $m_i \rightarrow \infty$, is originated actually by the finiteness of m_i . Generally, the infinite m_i limit is taken within the conception that finite m_i corrections would be of the order of the electron-to-ion mass ratio m_e/m_i , or perhaps $(m_e/m_i)^{1/2}$, negligible in a certain approximation; the correction found here for finite m_i does not go to zero as $m_i \rightarrow \infty$. This apparent paradox can be explained as follows: In order to avoid transient effects one assumes generally that the field was switched on at time τ , very far in the remote past ($\tau \rightarrow -\infty$). If we first let $m_i \rightarrow \infty$, the ions will not respond to the field no matter how long we wait (i.e., even when we let $\tau \rightarrow -\infty$), as discussed in the preceding paragraph. On the other hand, if m_i is first taken to be finite, the ions will move under the action of the field and after a long time some sort of non-thermal, oscillating equilibrium will be reached; if now we let $m_i \rightarrow \infty$, some static field-dependent part of the ion-ion correlations will not go to zero, not being dependent on m_i (as the thermal correlations are not).⁷

The time required for the ion-ion correlations to reach equilibrium is an average of the ion Landau damping time $\bar{\gamma}_i^{-1}$. We take the limit $|\tau| \bar{\gamma}_i \rightarrow \infty$ and then the limit $m_i \rightarrow \infty$ ($\bar{\gamma}_i$ depends on m_i and goes to zero

as $m_i \rightarrow \infty$.) For times such that $|\tau|\bar{\gamma}_i = \mathcal{O}(1)$, complicated ion transient effects would enter the picture, while for $|\tau|\bar{\gamma}_i < \mathcal{O}(1)$ the formulation of Ref. 6 would seem to be valid. In the last case, some questions could arise on whether other transient effects would have decayed off so early.

There is a second requirement in the Dawson-Oberman model, and that is the selection of a zero-order electron distribution function. This shortcoming will not be eliminated here. However, this problem is physically much less interesting than that connected with the selection of an ion-ion correlation function. Moreover, for moderate fields, to be considered in detail in the last section, the non-thermal modification of the ion correlations affect much more the nonlinear results than the deviation of the electron distribution function from its Maxwellian value.

Finally we point out that the equations to be considered are formally solved for an arbitrary field intensity. However, some physical assumptions are introduced; the restrictions that they may impose on the field intensity are discussed in the last section.

II. THE NONLINEAR CURRENT DENSITY

We consider an infinite, homogeneous plasma composed of electrons and one species of ions, inside which there is a spatially uniform, oscillating electric field. (For waves of long but finite wavelength, the uniform field assumption is equivalent to the dipole approximation.) From the Liouville equation, the usual BBGKY hierarchy of equations⁸ is derived. The equation for the one-particle distribution function of the α species, f_α , reads

$$\frac{\partial f_{\alpha}^1}{\partial t} + \frac{q_{\alpha}}{m_{\alpha}} \underline{\underline{E}} \sin \omega t \cdot \frac{\partial f_{\alpha}^1}{\partial \underline{\underline{v}}_1} = \frac{q_{\alpha}}{m_{\alpha}} \frac{\partial}{\partial \underline{\underline{v}}_1} \cdot \int d\underline{\underline{r}}_2 d\underline{\underline{v}}_2 \frac{\partial \phi^{12}}{\partial \underline{\underline{r}}_1} \sum_{\beta} q_{\beta} n_{\beta} g_{\alpha\beta}^{12} \quad (1)$$

where $\underline{\underline{E}} \sin \omega t$ is the field prevailing inside the plasma, $\phi^{12} = |\underline{\underline{r}}_1 - \underline{\underline{r}}_2|^{-1}$, and $g_{\alpha\beta}^{12}$ is the two-particle, phase-space correlation function between the α and β species (both α and β stand for electrons, e, or ions, i). The superscripts are particle indices.

Usually a kinetic equation for f_{α}^1 is derived⁸ by solving the equation for $g_{\alpha\beta}^{12}$ [see Eqs. (14) and (25) later] assuming that f_{α}^1 and f_{β}^2 remain frozen on the time scale during which $g_{\alpha\beta}^{12}$ changes and introducing this value of $g_{\alpha\beta}^{12}(f_{\alpha}^1, f_{\beta}^2)$ into (1); that assumption is certainly not satisfied when $\omega/\omega_p \geq 0(1)$.¹ On the other hand, ω is then large compared with the collision frequency, ν : for most fully ionized plasmas $\omega_p/\nu = 0(N_D)$ where N_D , the number of particles in a Debye sphere, is a large number. The first and third terms in (1) are of order of ωf_{α}^1 and νf_{α}^1 , respectively, so that to lowest order in ν/ω the driving field is balanced by the inertial force; we may then solve (1) by expanding f_{α}^1 in powers of ν/ω whenever $\nu/\omega \ll 1$. [We point out here that $\omega_p/\nu = 0(N_D)$ is only valid, a priori, for a weak field, E; for a strong field ν will depend on E. At this point, we simply assume ω and E such that $\nu(E)/\omega < 0(1)$, and leave the matter until the restrictions on the field intensity are discussed in the last section.]

Writing

$$f_{\alpha}^1 = f_{\alpha 0}^1 + \nu/\omega f_{\alpha 1}^1 + \dots \quad (2)$$

we obtain from (1) the zero-order equation

$$\frac{\partial f_{\alpha 0}^1}{\partial t} + \frac{q_{\alpha}}{m_{\alpha}} \frac{E}{\omega} \sin \omega t \cdot \frac{\partial f_{\alpha 0}^1}{\partial v_1} = 0 \quad (3)$$

which describes a purely reactive response to the field; the solution of

Eq. (3) is an arbitrary function of $u_1^{\alpha} \equiv v_1 + \omega \frac{E}{\omega} \cos \omega t$, where

$\frac{E}{\omega} = q_{\alpha}/m_{\alpha} \frac{E}{\omega^2}$. We normalize $f_{\alpha 0}^1$: $\int f_{\alpha 0}^1 dv_1 = 1$.

Consider now the equation for $f_{\alpha 1}^1$. It is a straightforward matter to find out that the naive expansion (2) breaks down for long times because of secular behavior: there is, for instance, a slow heating of the plasma and $\int f_{\alpha 0}^1 v_1^2 dv_1 \sim \nu t$. (This problem does not exist in the linear limit because for small E, the heating vanishes like E^2 .) The relaxation frequency for the ion-ion correlations is some average of the ion Landau damping rate; when this is of the order of or smaller than ν we will run (for the long-time limit stipulated at the end of Sec. I) into times such that the accumulated heating can be neglected no more. To include such a case in our treatment, the heating has to be taken into account and the mathematical secularities avoided. This is done by using the known method of multiple scales.⁹ We introduce a second time scale so that $f_{\alpha 0}^1$ can be written as $f_{\alpha 0}^1 = f_{\alpha 0}^1(v_1 + \omega \frac{E}{\omega} \cos \omega t, t^* \omega/\nu)$, $dt^*/dt = 1$. Now $f_{\alpha 0}^1$ satisfies (3)

in the fast scale only and there is a slow variation of $f_{\alpha 0}^1$ which appears in the next-order equation. The slow scale is introduced, of course, into the other terms of (2) too, but since we will not go beyond first-order terms, this is irrelevant here. The equation of order ν/ω obtained from (1), for $\alpha = e$, reads now:

$$\begin{aligned} \nu/\omega \left[\frac{\partial f_{e0}^1}{\partial(\nu/\omega t^*)} + \frac{\partial f_{e1}^1}{\partial t} + \frac{q_e}{m_e} \underline{E} \sin \omega t \cdot \frac{\partial f_{e1}^1}{\partial \underline{v}_1} \right] \\ = \frac{q_e n_e}{m_e} \frac{\partial}{\partial \underline{v}_1} \cdot \int d\underline{r}_2 d\underline{v}_2 \frac{\partial \phi}{\partial \underline{r}_1} (g_{ee}^{12} - g_{ei}^{12}) \end{aligned} \quad (4)$$

where use has been made of the neutrality condition, $q_e n_e + q_i n_i = 0$. We integrate over t in (4) and obtain

$$f_{e1}^1 \sim - \int dt \left[\frac{\partial f_{e0}^1}{\partial(\nu/\omega t^*)} + \frac{q_e}{m_e} \underline{E} \sin \omega t \cdot \frac{\partial f_{e1}^1}{\partial \underline{v}_1} - \frac{\omega}{\nu} C \right] \quad (5)$$

where C is the right-hand side of (4). (We point out here that the expansion $g_{\alpha\beta}^{12} = g_{\alpha\beta(0)}^{12} + \nu/\omega g_{\alpha\beta(1)}^{12} + \dots$ should have been made and only $g_{e\beta(0)}^{12}$ should appear in C ; $g_{\alpha\beta(0)}^{12}$ is to be found later from the equation for $g_{\alpha\beta}^{12}$ where $f_{\alpha 0}^1$ and $f_{\beta 0}^2$ will substitute for f_{α}^1 and f_{β}^2 . Since we will not go beyond the zero-order term in $g_{\alpha\beta}^{12}$ we write throughout the paper $g_{\alpha\beta}^{12}$ for $g_{\alpha\beta(0)}^{12}$.)

The introduction of a new time scale has given us an additional freedom; this has to be used to avoid secularities in the expansion. We

want f_{el}^1 to be bounded in time, If the integrand in (5) were of fixed sign for large t , $\lim_{t \rightarrow \infty} f_{el}^1$ could be finite only if that integrand would vanish in the limit $t \rightarrow \infty$. Thus, an equation for the slow variation of f_{eo}^1 would be obtained. However, if some terms inside the bracket in (5) are oscillating, as will be actually the case, the separation of the secular terms cannot be made until the form of C is known in detail. (The analysis would seem to be simpler for any particular moment of f_{eo}^1 than for f_{eo}^1 itself, but in any event a knowledge of C is required.)

The slow variation of f_{eo}^1 , therefore, can be determined only after both g_{ei}^{12} and g_{ee}^{12} are known. The equation for g_{ee}^{12} in the presence of a strong, high-frequency field is complicated and does not become simplified in the limit $m_i \rightarrow \infty$, as is the case for g_{ei} and g_{ii} , which will be studied in Secs. III and IV. Thus, f_{eo}^1 will remain unknown and our results will be given for an arbitrary f_{eo}^1 . On the other hand, something important can be established about the slow-variable dependence of f_{eo}^1 : it can only involve even powers of \underline{E} . In effect, along an axis aligned with the field there is no preferential direction for the long times covering many periods of the field because $\underline{E} \sin \omega t$ changes sign every half-period; it is quite apparent that if f_{eo}^1 is even in \underline{u}^e , at a given time, it will remain so throughout the slow-scale development. After transient effects in the fast scale have disappeared, we should take f_{eo}^1 to be an even function of \underline{u}_1^e . The definite selection made in Ref. 6 of a Maxwellian distribution is arbitrary, although quite probably the correct results for the current would differ only by simple numerical factors. In the linear limit, of course, we can take

f_{e0}^1 to be Maxwellian because corrections would be of the order of E^2 .

Going back to Eq. (4), we multiply the equation by $n_e q_e v_1$ and integrate over v_1 . There results

$$\frac{\nu}{\omega} \frac{\partial j_{\omega 0}}{\partial(\nu/\omega t^*)} + \frac{\partial j_{\omega 1}}{\partial t} = \frac{q_e^3 n_e^2}{m_e} \int dr_{\omega 2} \frac{\partial \phi_{\omega 12}}{\partial r_{\omega 12}} \int g_{ei}^{12} dv_{\omega 1} dv_{\omega 2} \quad (7)$$

where $j_{\omega 0} = n_e q_e \int f_{e0}^1 v_1 dv_{\omega 1}$ and $j_{\omega 1} = n_e q_e \int \nu/\omega f_{e1}^1 v_1 dv_{\omega 1}$. Obviously there is no contribution to (7) from g_{ee}^{12} ; as for the last term on the left-hand side of (4) we point out that our normalization for f_{e0}^1 implies

$$\int f_{e1}^1 dv_{\omega 1} = 0. \text{ We can write}$$

$$j_{\omega 0} = n_e q_e \int f_{e0}^1 \left[u_{\omega 1}^e - \omega \underline{\underline{\epsilon}}_e \cos \omega t \right] du_{\omega 1}^e = j_{\omega 0}^* - n_e q_e \omega \underline{\underline{\epsilon}}_e \cos \omega t.$$

Thus, $\partial j_{\omega 0} / \partial(\nu/\omega t^*) = \partial j_{\omega 0}^* / \partial(\nu/\omega t^*)$; since f_{e0}^1 is even in $u_{\omega 1}^e$, $j_{\omega 0}^* = 0$ and we can drop the first term in (7). The zero-order current $j_{\omega 0}$ is purely reactive, then, and satisfies the equation

$$\frac{\partial j_{\omega 0}}{\partial t} - \frac{q_e}{m_e} \underline{\underline{\epsilon}}_e \sin \omega t = 0$$

obtained from (3) for $\alpha = e$.

Equation (7) reduces to

$$\frac{\partial j_{\omega 1}}{\partial t} = - \frac{q_e^3 n_e^2}{m_e} (2\pi)^{-3} \int ik \underline{\underline{\phi}} \int g_{ei}^{\omega 12} dv_{\omega 1} dv_{\omega 2} \quad (9)$$

where for later convenience we have made some rearrangement by

introducing the Fourier transforms

$$\tilde{\phi}(\underline{k}) = \int d\underline{r}_{12} \phi^{12} \exp[-i\underline{k} \cdot \underline{r}_{12}] = \frac{4\pi}{k^2}, \quad (10)$$

$$\tilde{g}_{ei}^{12}(\underline{k}, \underline{v}_1, \underline{v}_2) = \int d\underline{r}_{12} g_{ei}^{12}(\underline{r}_{12}, \underline{v}_1, \underline{v}_2) \exp[-i\underline{k} \cdot \underline{r}_{12}], \quad (11)$$

where we have taken advantage of the spatial homogeneity. In the next two sections we shall obtain an expression for the right-hand side of (9) in terms of f_{eo}^1 .

Finally, let us multiply (4) by $n_e m_e v_1^2/2$ and integrate over \underline{v}_1 and t .

We obtain

$$U_1 \sim - \int^t dt \left[\frac{\nu}{\omega} \frac{\partial U_o}{\partial(\nu/\omega t^*)} - \underline{j}_1 \cdot \underline{E} \sin \omega t - n_e \frac{m_e}{2} \int C v_1^2 d\underline{v}_1 \right] \quad (12)$$

where $U_o = 1/2 n_e m_e \int f_{eo}^1 v_1^2 d\underline{v}_1$ and $U_1 = 1/2 n_e m_e \int \nu/\omega f_{e1}^1 v_1^2 d\underline{v}_1$.

It follows immediately that

$$\frac{\partial U_o}{\partial(\nu/\omega t^*)} = \frac{\partial}{\partial(\nu/\omega t^*)} 1/2 n_e m_e \int f_{eo}^1 \left[\frac{u_e}{v_1} \right]^2 d\underline{v}_1$$

so that the first term inside the bracket of (12) does not depend on t . The third term involves only g_{ei}^{12} because there can be no self-heating of the electrons. In the limit $m_i \rightarrow \infty$, both \underline{j}_1 and $\int C v_1^2 d\underline{v}_1$ will come out to be purely oscillating in t (there are additional dependences on $\nu/\omega t^*$). To avoid secularities, therefore,

$$\frac{\partial U_o}{\partial t^*} = \langle \underline{j}_1 \cdot \underline{E} \sin \omega t \rangle \quad (13)$$

in the limit $t \rightarrow \infty$, where $\langle \rangle$ indicates averaging over the fast variable; the average does not vanish because of those terms in j_1 that vary as $\sin \omega t$. Equation (13) describes the slow heating of the electrons. We point out that if f_{e0}^1 were Maxwellian, there would be only one unknown parameter, the temperature T_e . Hence, (13) would describe completely the evolution of f_{e0}^1 , i. e., the slow nonlinear change in T_e (j_1 being a nonlinear function of T_e).

III. THE ELECTRON-ION CORRELATION FUNCTION

The equation for the electron-ion correlation function, g_{ei} , reads⁸

$$\begin{aligned} & \left[\frac{\partial}{\partial t} + (\underline{v}_1 - \underline{v}_2) \cdot \frac{\partial}{\partial \underline{r}_{12}} + \underline{E} \sin \omega t \cdot \left(\frac{q_e}{m_e} \frac{\partial}{\partial \underline{v}_1} + \frac{q_i}{m_i} \frac{\partial}{\partial \underline{v}_2} \right) \right] g_{ei}^{12} \\ &= q_e q_i \frac{\partial \phi^{12}}{\partial \underline{r}_{12}} \cdot \left(\frac{1}{m_e} \frac{\partial}{\partial \underline{v}_1} - \frac{1}{m_i} \frac{\partial}{\partial \underline{v}_2} \right) f_e^1 f_e^2 + \frac{q_e}{m_e} \frac{\partial f_e^1}{\partial \underline{v}_1} \cdot \frac{\partial}{\partial \underline{r}_{12}} \int d\underline{r}_3 \\ & \quad \times \phi^{13} \int d\underline{v}_3 \sum_{\alpha} q_{\alpha}^n g_{i\alpha}^{23} - \frac{q_i}{m_i} \frac{\partial f_i^2}{\partial \underline{v}_2} \cdot \frac{\partial}{\partial \underline{r}_{12}} \int d\underline{r}_3 \phi^{23} \int d\underline{v}_3 \sum_{\alpha} q_{\alpha}^n g_{e\alpha}^{13} . \end{aligned} \quad (14)$$

Again the homogeneity of the plasma has been used, by writing $g_{\alpha\beta}^{mn}(\underline{r}_{\underline{m}}, \underline{r}_{\underline{n}}) = g_{\alpha\beta}^{mn}(\underline{r}_{\underline{m}} - \underline{r}_{\underline{n}})$. To obtain (14) we have dropped two terms from the exact BBGKY equation, as is usually done in plasma kinetic theory. One is $\partial \phi^{12} / \partial \underline{r}_{12} \cdot (\partial / m_e \partial \underline{v}_1 - \partial / m_i \partial \underline{v}_2) g_{ei}^{12}$; the other involves the three-particle correlation function $h_{ei\alpha}^{123}$. (Similar terms will be dropped from the BBGKY equation for g_{ii} in the next section.) In general, those terms

can be neglected whenever $|g_{ei}/f_e f_i| \ll 1$ and $|h_{ei\alpha}| \ll f_e |g_{i\alpha}|$, $f_i |g_{e\alpha}|$, respectively. [The first term can be also dropped on the basis of a comparison to the second term on the left side of (14); their ratio is of the order of $e^2/|r_{12}| \kappa T$. Thus, (14) can be valid only for $|r_{12}| \gg e^2/\kappa T$.

The exclusion of the small impact region is common and amounts to a weak indeterminacy in a cutoff to be introduced in the subsequent computation of some integrals.] Not far from thermal equilibrium, a comparison with known results provides confirmation of the previously mentioned inequalities.

For the arbitrary field intensity we are considering, we now make the ansatz that the terms dropped are negligible and later, in the last section, discuss the restrictions that this may impose on the field intensity.

To obtain a zero-order result for g_{ei}^{12} we can substitute $f_{\alpha 0}^m$ for f_{α}^m in (14). Next we introduce the transformation

$$t = t, \quad \underline{u}_m^\alpha = \underline{v}_m + \omega \underline{\varepsilon}_\alpha \cos \omega t,$$

$$\underline{\rho}_m^\alpha = \underline{r}_m + \underline{\varepsilon}_\alpha \sin \omega t$$

so that $\underline{\rho}_{12}^e \equiv \underline{\rho}_1^e - \underline{\rho}_2^i = \underline{r}_{12} + \underline{\varepsilon}_{ei} \sin \omega t$, ($\underline{\varepsilon}_{\alpha\beta} = \underline{\varepsilon}_\alpha - \underline{\varepsilon}_\beta$). The equation for g_{ei}^{12} becomes

$$\begin{aligned} & \left[\frac{\partial}{\partial t} + (\underline{u}_1^e - \underline{u}_2^i) \cdot \frac{\partial}{\partial \underline{\rho}_{12}} \right] g_{ei}^{12} = q_e q_i \frac{\partial \phi^{12}}{\partial \underline{\rho}_{12}} \cdot \left(\frac{1}{m_e} \frac{\partial}{\partial \underline{u}_1^e} - \frac{1}{m_i} \frac{\partial}{\partial \underline{u}_2^i} \right) \\ & \times f_{eo}^1 f_{io}^2 + \frac{q_e q_i^n}{m_e} \frac{\partial f_{eo}^1}{\partial \underline{u}_1^e} \cdot \frac{\partial}{\partial \underline{\rho}_{12}} \int d\underline{r}_3 \phi^{13} \left(\int g_{ie}^{23} d\underline{u}_3^e - \int g_{ii}^{23} d\underline{u}_3^i \right) \\ & - \frac{q_e q_i^n}{m_i} \frac{\partial f_{io}^2}{\partial \underline{u}_2^i} \cdot \frac{\partial}{\partial \underline{\rho}_{12}} \int d\underline{r}_3 \phi^{23} \left(\int g_{ee}^{13} d\underline{u}_3^e - \int g_{ei}^{13} d\underline{u}_3^i \right). \end{aligned} \quad (15)$$

We integrate (15) over \underline{u}_2^i to obtain

$$\left[\frac{\partial}{\partial t} + \underline{u}_1^e \cdot \frac{\partial}{\partial \underline{\rho}_{12}} \right] G_{ei}^{12} = \frac{q_e q_i}{m_e} \frac{\partial \phi^{12}}{\partial \underline{\rho}_{12}} \cdot \frac{\partial f_{eo}^1}{\partial \underline{u}_1^e} + \frac{q_e^2 n_e}{m_e} \frac{\partial f_{eo}^1}{\partial \underline{u}_1^e} \frac{\partial}{\partial \underline{\rho}_{12}}$$

$$\times \int d\underline{r}_3 \phi^{13} \left(\int G_{ei}^{23} d\underline{u}_3^e - \int g_{i1}^{32} d\underline{u}_i^2 d\underline{u}_i^3 \right) - \frac{\partial}{\partial \underline{\rho}_{12}} \cdot \int \underline{u}_2^i g_{ei}^{12} d\underline{u}_2^i \quad (16)$$

where $G_{ei}^{mn} = \int g_{ei}^{mn} d\underline{u}_i^n$ and the equality $g_{\alpha\beta}^{mn} = g_{\beta\alpha}^{nm}$ has been used. Those terms in (15) proportional to $\partial f_{io}^2 / \partial \underline{u}_2^i$ do not contribute to (16) since $f_{io}^2 (\underline{u}_2^i \rightarrow \pm\infty) \rightarrow 0$, as required from any distribution function.

While we would have to solve an integral equation to obtain g_{ei}^{12} from (15), it is very simple to obtain G_{ei}^{12} from (16); of course, the last term of (16) is quite awkward but we expect that it will drop out in the infinite ion mass limit, which ultimately we shall take. We also observe that \tilde{g}_{ee} has disappeared from our equation; it will be considered no more in this paper.

Given any function of \underline{r}_{12} , $A(\underline{r}_{12})$, we can introduce the following Fourier transforms: $\tilde{A}(\underline{k}) = \int \exp[-i\underline{k} \cdot \underline{r}_{12}] d\underline{r}_{12} A(\underline{r}_{12})$ [as in Eqs. (10) and (11)] and

$$\tilde{\tilde{A}}(\underline{k}) \equiv \int \exp[-i\underline{k} \cdot \underline{\rho}_{12}] d\underline{\rho}_{12} A(\underline{\rho}_{12}) = \exp[-i\underline{k} \cdot \underline{\varepsilon}_{ei} \sin \omega t]$$

$$\times \int \exp[-i\underline{k} \cdot \underline{r}_{12}] d\underline{r}_{12} A(\underline{r}_{12}) = \exp[-i\underline{k} \cdot \underline{\varepsilon}_{ei} \sin \omega t] \tilde{A}(\underline{k}). \quad (17)$$

Then, making a Fourier analysis of Eq. (16) with respect to the variable $\underline{\rho}_{12}$ yields

$$\begin{aligned}
 \left(\frac{\partial}{\partial t} + \mathbf{ik} \cdot \mathbf{u}_1^e \right) G_{ei}^{12} &= \frac{q_e q_i}{m_e} \mathbf{ik} \cdot \frac{\partial f_{eo}^1}{\partial \mathbf{u}_1^e} \tilde{\phi} \exp \left[-\mathbf{ik} \cdot \boldsymbol{\varepsilon}_{ei} \sin \omega t \right] \\
 &+ \frac{q_e^2 n_e}{m_e} \mathbf{ik} \cdot \frac{\partial f_{eo}^1}{\partial \mathbf{u}_1^e} \frac{4\pi}{k} \left(\int G_{ei}^{32} d\mathbf{u}_3^e - \exp \left[-\mathbf{ik} \cdot \boldsymbol{\varepsilon}_{ei} \sin \omega t \right] \right) \\
 &\times \int g_{ii}^{32} d\mathbf{u}_3^i d\mathbf{u}_2^i - \mathbf{ik} \cdot \int \mathbf{u}_2^i g_{ei}^{12} d\mathbf{u}_2^i \quad (18)
 \end{aligned}$$

where use has been made of (10) and (17).

We make the ansatz $\tilde{g}_{ii}^{12} = \sum_p \tilde{g}_{ii}^{p} (1, 2) e^{-ip\omega t}$ and define

$\tilde{Q}_{ii}^p = \int \tilde{g}_{ii}^p (m, n) d\mathbf{v}_m d\mathbf{v}_n$. Next, we solve (18) under the assumption of vanishing perturbations in the remote past; a formal integration of (18) yields

$$\begin{aligned}
 \tilde{G}_{ei}^{12} &= \frac{\mathbf{ik}}{m_e} \cdot \frac{\partial f_{eo}^1}{\partial \mathbf{u}_1^e} \int_{-\infty}^t d\tau \exp \left[-\mathbf{ik} \cdot \mathbf{u}_1^e (t-\tau) \right] \exp \left[-\mathbf{ik} \cdot \boldsymbol{\varepsilon}_{ei} \sin \omega \tau \right] \\
 &\times \left(q_e q_i \tilde{\phi} - q_e^2 \frac{4\pi}{k} n_e \sum_p \tilde{Q}_{ii}^p e^{-ip\omega \tau} \right) + \frac{\mathbf{ik}}{m_e} \cdot \frac{\partial f_{eo}^1}{\partial \mathbf{u}_1^e} q_e^2 n_e \\
 &\times \frac{4\pi}{k} \int_{-\infty}^t d\tau \exp \left[-\mathbf{ik} \cdot \mathbf{u}_1^e (t-\tau) \right] \int G_{ei}^{32} (\tau) d\mathbf{u}_3^e \\
 &- \mathbf{ik} \cdot \int \mathbf{u}_2^i d\mathbf{u}_2^i \int_{-\infty}^t d\tau \exp \left[-\mathbf{ik} \cdot \mathbf{u}_1^e (t-\tau) \right] \tilde{g}_{ei}^{12} (\tau).
 \end{aligned}$$

If this equation is integrated over \mathbf{u}_1^e we obtain the following integral

equation for $\tilde{Q}_{ei}^p \equiv \int \tilde{G}_{ei}^p d\mathbf{u}_1^e$:

$$\begin{aligned}
 \tilde{Q}_{ei}^{\approx}(t) &= \frac{\omega_p^2}{k^2} \int du_1^e ik \cdot \frac{\partial f_{e0}^1}{\partial u_1^e} \int_{-\infty}^t d\tau \tilde{Q}_{ei}^{\approx}(\tau) \exp[-ik \cdot u_1^e(t-\tau)] \\
 &= -\frac{\omega_p^2}{k^2} \int du_1^e ik \cdot \frac{\partial f_{e0}^1}{\partial u_1^e} \int_{-\infty}^t d\tau \exp[-ik \cdot u_1^e(t-\tau)] \exp[-ik \cdot \varepsilon_{ei} \sin \omega t] \\
 &\quad \times \left(\frac{1}{n_i} + \sum_p \tilde{Q}_{ii}^p \right) e^{-ip\omega\tau} \int_{-\infty}^t d\tau \int \tilde{g}_{ei}^{12}(\tau) \exp[-ik \cdot u_1^e(t-\tau)] ik \cdot u_2^i du_2^i du_1^e .
 \end{aligned} \tag{19}$$

The solution to (19) can be written as

$$\tilde{Q}_{ei}^{\approx} = Q_1 + Q_2 \equiv \sum_p e^{-ip\omega t} \left[\frac{J_p(k \cdot \varepsilon_{ei}) \chi_{-p}}{n_i D_{-p}} (1 + n_i \tilde{Q}_{ii}^o) + B_p \right] + Q_2 \tag{20}$$

where Q_1 and Q_2 are due to the first and second terms on the right-hand

side of (19), respectively. To find Q_1 we have used the identity $\exp[-ik \cdot \varepsilon_{ei} \sin \omega t] =$

$$\sum_p J_p(k \cdot \varepsilon_{ei}) e^{-ip\omega t} . \text{ In (20),}$$

$$D_{-p} \equiv 1 + \chi_{-p} = 1 + \frac{\omega_p^2}{k^2} \int \frac{k \cdot \frac{\partial f_{e0}^1}{\partial u_1^e} du_1^e}{p\omega - k \cdot u_1^e + i\varepsilon} ;$$

$$B_p = \frac{\chi_{-p}}{D_{-p}} \sum_{-p, s \neq 0} J_{p-s} \tilde{Q}_{ii}^s . \tag{21}$$

The plasma frequency is $\omega_p = (4\pi q_e^2 n_e / m_e)^{1/2}$, and $\varepsilon > 0$ and goes to zero.

If the limit $m_i \rightarrow \infty$ we expect Q_{ii}^s ($s \neq 0$) and $\int g_{ei}^{12} \frac{u_i}{m_2} du_2^i$ to vanish; if this is the case, B_p and Q_2 will vanish and

$$\tilde{Q}_{ei}^0 = \sum_p e^{-ip\omega t} \frac{J_p \chi_{-p}}{n_i D_{-p}} (1 + n_i \tilde{Q}_{ii}^0) \quad (22)$$

so that using (17), Eq. (9) reduces to

$$\begin{aligned} \frac{\partial j_1}{\partial t} &= \frac{-q_e^3 n_e^2}{2\pi m_e} \int \frac{ik dk}{k^2} \exp\left[ik \cdot \varepsilon_{ei} \sin \omega t \right] \sum_p e^{-ip\omega t} \frac{J_p \chi_{-p}}{n_i D_{-p}} (1 + n_i \tilde{Q}_{ii}^0) \\ &= - \frac{ie^3 n_e Z}{2\pi m_e} \int \frac{k dk}{k^2} \sum_\ell e^{i\ell \omega t} \sum_p \frac{J_p J_{p+\ell}}{D_{-p}} (1 + n_i \tilde{Q}_{ii}^0) \end{aligned} \quad (23)$$

If we write above $\tilde{Q}_{ii}^0 = 0$, we recover the result found in Ref. 6 by using the Dawson-Oberman model. The proper value to be used for \tilde{Q}_{ii}^0 will be obtained in the following section.

Also, if the limit $E \rightarrow 0$ is taken in (23) and terms linear in E are retained, the results of Ref. 1 and Refs. 2 and 3 are recovered by choosing

$\tilde{Q}_{ii}^0 = 0$ and

$$\tilde{Q}_{ii}^0 = - \frac{1}{n_i} \frac{k_i^2}{k_i^2 + k_e^2 + k^2} \quad (24)$$

respectively. Equation (24) is the known expression for the thermal ion-ion correlations, where $k_\alpha^2 = 4\pi q_\alpha^2 n_\alpha (\kappa T_\alpha)^{-1}$, κ being Boltzman's constant.

IV. THE ION-ION CORRELATION FUNCTION

We now proceed to determine \tilde{Q}_{ii}^0 and to show that \tilde{Q}_{ii}^s ($s \neq 0$) and $\int g_{ei}^{\approx 12} u_2^i du_2^i$ vanish in the limit $m_i \rightarrow \infty$. The equation for g_{ii}^{12} reads

$$\begin{aligned} & \left[\frac{\partial}{\partial t} + (\underline{v}_1 - \underline{v}_2) \cdot \frac{\partial}{\partial \underline{r}_{12}} + \frac{q_i}{m_i} E \sin \omega t \cdot \left(\frac{\partial}{\partial \underline{v}_1} + \frac{\partial}{\partial \underline{v}_2} \right) \right] g_{ii}^{12} \\ &= \frac{q_i^2}{m_i} \frac{\partial \phi^{12}}{\partial \underline{r}_{12}} \cdot \left(\frac{\partial}{\partial \underline{v}_1} - \frac{\partial}{\partial \underline{v}_2} \right) f_{io}^1 f_{io}^2 + \frac{q_i q_e n_e}{m_i} \frac{\partial f_{io}^1}{\partial \underline{v}_1} \cdot \frac{\partial}{\partial \underline{r}_{12}} \int d\underline{r}_3 \phi^{13} \int d\underline{v}_3 g_{ie}^{23} \\ & - \int d\underline{v}_3 g_{ii}^{23} - \frac{q_i q_e n_e}{m_i} \frac{\partial f_{io}^2}{\partial \underline{v}_2} \cdot \frac{\partial}{\partial \underline{r}_{12}} \int d\underline{r}_3 \phi^{23} \left(\int d\underline{v}_3 g_{ie}^{13} - \int d\underline{v}_3 g_{ii}^{13} \right) \end{aligned} \quad (25)$$

We introduce the transformation $t = t$, $u_n^\alpha = \underline{v}_n + \omega \underline{\varepsilon}_n \cos \omega t$, and Fourier analyze with respect to \underline{r}_{12} to obtain

$$\begin{aligned} \left[\frac{\partial}{\partial t} + i\mathbf{k} \cdot (\underline{u}_1^i - \underline{u}_2^i) \right] g_{ii}^{\approx 12} &= \frac{q_i^2}{m_i} i\mathbf{k} \cdot \tilde{\phi} \cdot \left(\frac{\partial}{\partial \underline{u}_1^i} - \frac{\partial}{\partial \underline{u}_2^i} \right) f_{io}^1 f_{io}^2 \\ &+ \frac{q_i q_e n_e}{m_i} i\mathbf{k} \cdot \frac{\partial f_{io}^1}{\partial \underline{u}_1^i} \frac{4\pi}{k^2} \left\{ \exp[i\mathbf{k} \cdot \underline{\varepsilon}_{ei} \sin \omega t] \int g_{ei}^{\approx 32} du_3^e - \int g_{ii}^{\approx 32} du_i^3 \right\} \\ &- \frac{q_i q_e n_e}{m_i} i\mathbf{k} \cdot \frac{\partial f_{io}^2}{\partial \underline{u}_2^i} \frac{4\pi}{k^2} \left\{ \exp[-i\mathbf{k} \cdot \underline{\varepsilon}_{ei} \sin \omega t] \int g_{ie}^{\approx 13} du_3^e - \int g_{ii}^{\approx 13} du_i^3 \right\} \end{aligned} \quad (26)$$

where Eqs. (10) and (17) have been used.

Now we complete the ansatz made in Sec. III, $\tilde{g}_{ii}^{mn} = \sum_s e^{-is\omega t} \tilde{g}_{ii}^{s(m,n)}$,

by assuming that for large m_i ,

$$f_{io}^i(u^i) = f_{im}^i(u^i) + o(1), \quad g_{ei}^{n2} = f_{io}^2 G_{ei}^{n2} + o(1),$$

$$g_{ii}^0 = f_{io} f_{io} \tilde{Q}_{ii}^0 + o(1) \quad (27)$$

where f_{im}^i is a Maxwellian distribution; the symbol $o(1)$ represents a quantity that goes to zero as $m_i \rightarrow \infty$. As obtained from (26), the equation for \tilde{g}_{ii}^s ($s \neq 0$) reads for very large m_i ,

$$\begin{aligned} [-s\omega + k \cdot (u_1^i - u_2^i)] \tilde{g}_{ii}^s(1,2) &\approx \frac{q_e q_i^n e}{k T_i} \frac{4\pi}{k^2} k \cdot \left[u_1^i f_{im}^1 \int \tilde{g}_{ii}^s(3,2) du_3^i \right. \\ &\quad \left. - u_2^i f_{im}^2 \int \tilde{g}_{ii}^s(1,3) du_3^i \right] - \frac{q_e q_i^n e}{m_i} \frac{4\pi}{k} k \\ &\times \left[u_1^i f_{im}^1 f_{im}^2 \sum_{\ell} J_{\ell} \tilde{Q}_{ei}^{s+\ell} - u_2^i f_{im}^2 f_{im}^1 \sum_{\ell} J_{\ell} \tilde{Q}_{ei}^{s-\ell}(-k) \right] \quad (28) \end{aligned}$$

where use has been made of the identity $\tilde{Q}_{ei}^{s}(-k) = \tilde{Q}_{ie}^{s}(k)$. We have called \tilde{Q}_{ei}^s the bracket multiplying $e^{-is\omega t}$ in Eq. (20) and dropped Q_2 because with the ansatz (27), $\int g_{ei}^{12} u_2^i du_2^i \equiv 0$.

If the limit $m_i \rightarrow \infty$ is taken in (28), $f_{im}^i \rightarrow \delta(u^i)$ and since $u^i \delta(u^i) \equiv 0$,

there results

$$[-s\omega + k \cdot (u_1^i - u_2^i)] \tilde{g}_{ii}^s = 0,$$

so that $g_{ii}^s = 0$ for $s \neq 0$; in this limit, therefore, $B_p \rightarrow 0$, [see Eq. (21)].

As for \tilde{g}_{ii}^0 , we have from Eqs. (26) and (27):

$$\begin{aligned} \underline{k} \cdot (\underline{u}_1^i - \underline{u}_2^i) f_{im}^1 f_{im}^2 \tilde{Q}_{ii}^0 &\approx -\frac{q_i^2}{\kappa T_i} \tilde{\phi} \underline{k} \cdot (\underline{u}_1^i - \underline{u}_2^i) f_{im}^1 f_{im}^2 \\ &+ \frac{q_e q_i n_e}{\kappa T_i} \frac{4\pi}{k^2} \underline{k} \cdot (\underline{u}_1^i - \underline{u}_2^i) f_{im}^1 f_{im}^2 \tilde{Q}_{ii}^0 - \frac{q_e q_i n_e}{\kappa T_i} \frac{4\pi}{k^2} \\ &\times f_{im}^1 f_{im}^2 \underline{k} \cdot \left[\underline{u}_1^i \sum_{\ell} J_{\ell} \tilde{Q}_{ei}^{\ell} - \underline{u}_2^i \sum_{\ell} J_{\ell} \tilde{Q}_{ei}^{-\ell}(-\underline{k}) \right] . \quad (29) \end{aligned}$$

We divide throughout by $(f_{im}^1 f_{im}^2)$ and take the limit $m_i \rightarrow \infty$; since $B_{\ell} \rightarrow 0$, $\tilde{Q}_{ei}^{\ell} \rightarrow J_{\ell} \chi_{-l} (n_i D_{i-l})^{-1} (1 + n_i \tilde{Q}_{ii}^0)$ and we obtain

$$\begin{aligned} \underline{k} \cdot (\underline{u}_1^i - \underline{u}_2^i) \left[\tilde{Q}_{ii}^0 + \frac{q_i^2}{\kappa T_i} \tilde{\phi} - \frac{q_e q_i n_e}{\kappa T_i} \frac{4\pi}{k^2} \tilde{Q}_{ii}^0 \right. \\ \left. + \frac{q_e q_i n_e}{\kappa T_i} \frac{4\pi}{k^2} \sum_{\ell} \frac{J_{\ell}^2 \chi_{-l}}{n_i D_{i-l}} (1 + n_i \tilde{Q}_{ii}^0) \right] = 0 . \quad (30) \end{aligned}$$

We have made use of the equality $\tilde{Q}_{ii}^0(-\underline{k}) = \tilde{Q}_{ii}^0(\underline{k})$, which follows from simple symmetry considerations about $\int g_{ii}^{12} dv_1 dv_2$; then

$$\begin{aligned} \sum_{\ell} J_{\ell}(\underline{k} \cdot \underline{\varepsilon}_{\underline{m}e}) \frac{J_{\ell}(\underline{k} \cdot \underline{\varepsilon}_{\underline{m}e}) \chi_{-l}(\underline{k})}{D_{-l}(\underline{k})} [1 + n_i \tilde{Q}_{ii}^0(\underline{k})] \\ = \sum_{\ell} J_{\ell}(\underline{k} \cdot \underline{\varepsilon}_{\underline{m}e}) \frac{J_{-\ell}(-\underline{k} \cdot \underline{\varepsilon}_{\underline{m}e}) \chi_{\ell}(-\underline{k})}{D_{\ell}(-\underline{k})} [1 + n_i \tilde{Q}_{ii}^0(-\underline{k})] \end{aligned}$$

and the brackets in (29) can be written: $(\underline{u}_1^i - \underline{u}_2^i) \sum_{\ell} J_{\ell} \tilde{Q}_{ei}^{\ell}$, as we have

done in (30) . Equation (30) yields immediately

$$\tilde{Q}_{ii}^o = - \frac{1}{n_i} \frac{k_i^2 \sum_{\ell} J_{\ell}^2 / D_{-\ell}}{k^2 + k_i^2 \sum_{\ell} J_{\ell}^2 / D_{-\ell}} \quad (31)$$

and

$$1 + n_i \tilde{Q}_{ii}^o = \frac{k^2}{k^2 + k_i^2 \sum_{\ell} J_{\ell}^2 / D_{-\ell}} \quad (32)$$

\tilde{Q}_{ii}^o is a function of ω through the dispersion functions $D_{-\ell}$.

We can go back now to Eq. (15) and observe that our ansatz $g_{ei} = f_{im} G_{ei}$ satisfies (15) in the limit $m_i \rightarrow \infty$, by simply noticing Eq. (16) and the relation $\underline{u}^i \delta(\underline{u}^i) \equiv 0$. We have, therefore, obtained a consistent solution to Eqs. (15) and (26) by assuming the ansatz (27). We point out here that this ansatz has put no restriction on the equation for g_{ee} ; for $m_i \rightarrow \infty$, the equations for g_{ei} and g_{ee} do not involve g_{ee} , the equation for which, on the other hand, involves G_{ei} (although not g_{ei} and g_{ii}).

V. DISCUSSION

We have found an expression for the ion-ion correlations, \tilde{Q}_{ii}^o . It is an even, real function of k as it should. In the limit $E \rightarrow 0$ (when f_{eo}^1 should become Maxwellian), Eq. (31) becomes

$$\tilde{Q}_{ii}^o = -\frac{1}{n_i} \frac{k_i^2}{k_i^2 + k_e^2 + k^2} + O(E^2) \quad (33)$$

so that the thermal correlations are recovered. The corrections for small E are $O(E^2)$ so that in the linear limit one can use the thermal value of \tilde{Q}_{ii}^o , as proposed in Sec. I. (That \tilde{Q}_{ii}^o had to be even in E , follows from symmetry considerations.) We also observe that for $E \rightarrow 0$,

$$\tilde{Q}_{ei} = \frac{1}{n_i} \frac{k_e^2}{k_e^2 + k_i^2 + k^2} + O(E) \quad ;$$

the first term is the expression for the electron-ion thermal correlations.

Of the two functions that the Dawson-Oberman model requires and that the model does not provide, f_{eo} and \tilde{Q}_{ii}^o , we only determine \tilde{Q}_{ii}^o ; to find f_{eo} an analysis of the equation of g_{ee} has to be made and this is left for future work. On the other hand, the physical phenomenon studied here (the finite modifications that the field produces in the ion correlations for very large ion mass) seems quite interesting by itself, while not much physics is involved in the determination of f_{eo} . Moreover, for moderate fields, studied in more detail later in this section, the indeterminacy in f_{eo} is unimportant, while the precise form of \tilde{Q}_{ii}^o is quite essential.

When Eq. (32) is introduced into Eq(23) an expression for the current density results. Carrying into Eq. (13) the component of \underline{j}_1 that varies as $\sin \omega t$ provides a description of the heating, or absorption by the plasma particles of the field energy. Also, the expression for the current can be used in Maxwell's equations to study the propagation of an electromagnetic wave (of very large wavelength, $\lambda \gg |\underline{\epsilon}_{ei}|$, according to our uniform field assumption); we should notice that all harmonics of the fundamental frequency appear in the current and therefore will appear in the wave too, as follows from Ampère's law. However, $|\underline{j}_1| \ll |\underline{j}_0|$ and so an iterative procedure can be used, as in Ref. 6, whenever a detailed study of the propagation is wanted.

Our treatment is formally valid for any field intensity. However, we have made some assumptions and we want to discuss now any physical restrictions that these assumptions may impose on the intensity of the field. First of all we assumed that a (quasi-steady) equilibrium would be reached, expecting that the transients would die off. This would not be the case, certainly, if the presence of the field makes the plasma unstable. It is known¹⁰ that for $\omega \approx \omega_p$ an instability sets in for relatively weak fields $[\omega |\underline{\epsilon}_{ee}| \ll (\kappa T_e / m_e)^{1/2}]$; for $\omega < \omega_p$ there are instabilities too, but the threshold is larger: $\omega |\underline{\epsilon}_{ee}| / (\kappa T_e m_e^{-1})^{1/2} = 0(1)$.¹¹ Therefore, when $\omega \lesssim \omega_p$ our results will be valid only if E is less than the appropriate threshold field for instability.

A second restriction may be imposed by the ansatz made at the beginning of Sec. III in connection with Eq. (14) (as well as Eq. (25)).

Dropping the term $\partial\phi^{12}/\partial r_{e12} \cdot (\partial/m_e \partial v_{e1} - \partial/m_i \partial v_{i2}) g_{ei}^{12}$ from the BBGKY equation for g_{ei} , as we did by writing Eq. (14), is valid if

$|g_{ei}| \ll f_e f_i$ since we retain a term, the first on the right side of Eq. (14), that would be much larger than the neglected one. A similar comment

could be made for Eq. (25). For our results to be consistent we should

require, then, that $|G_{ei}/f_e| \ll 1$ and $|Q_{ii}^0| \ll 1$ (for $|r_{12}| \gg e^2/kT$).

\tilde{G}_{ei} can be easily obtained from Eqs. (18) and (22). By Fourier inversion

of \tilde{Q}_{ii}^0 and \tilde{G}_{ei} the consistency could be checked. Unfortunately, the

Fourier inversion results in too-complicated integrals, especially for

G_{ei} . To clarify the validity of our dropping the aforementioned terms

we have to argue in a different way, as follows:

As noticed by simply examining Eqs. (14) and (15), the electric field nearly ceases to appear explicitly in the equation for g_{ei} , after a proper

transformation is made; this happens also in the g_{ii} equation in Sec. IV

(and obviously in the g_{ee} equation, not studied here). In the new variables

the field enters in the expression for $\phi^{12} = |r_{12}|^{-1} \equiv |\rho_{12} - \epsilon_{ei} \sin \omega t|^{-1}$;

this difference aside, Eq. (15) is the same as the one in the absence of the

field in the original variables. If ϕ^{12} were not time dependent in the new

frame of reference, our dropping $\partial\phi^{12}/\partial\rho_{12} \cdot (\partial/m_e \partial u_{e1}^e - \partial/m_i \partial u_{i2}^i) g_{ei}^{12}$

would seem to be valid because its ratio to a term retained in Eq. (15)

(the second on the left-hand side) would appear to be of order of

$e^2 / \kappa T |\rho_{12} - \epsilon_{ei} \sin \omega t|^{-1}$ and, therefore, small except in the region $|r_{12}| \equiv |\rho_{12} - \epsilon_{ei} \sin \omega t| = 0 (e^2 / \kappa T)$. Although in the ρ_{12} frame this region is not at the origin, it is small and our lack of knowledge of the proper equation for g_{ei} inside it would result in the usual unimportant weak indeterminacy in some integrals. For the actually time-dependent ϕ^{12} that we have, the argument obviously holds when $|\epsilon_{ei}|$ is small compared with the Debye length. If $|\epsilon_{ei}|$ is comparable to or much larger than the Debye length, the region where Eq. (15) is not valid covers a large, important part in the ρ_{12} space when a whole period of the field is considered. Nevertheless, we point out that at any given time the region of indeterminacy is still quite small, $0(e^2 / \kappa T)$; although this is certainly not conclusive, we are inclined to believe on this basis that the term we are discussing can be validly dropped from the BBGKY equation, even when $|\epsilon_{ei}|$ is comparable or large compared with the Debye length.

Whenever $|g_{\alpha\beta}| \ll f_{\alpha} f_{\beta}$, an argument can be given to justify our dropping also the three-particle correlation function, $h_{\alpha\beta\gamma}$, from the $g_{\alpha\beta}$ equation, as we did in Eqs. (14) and (25). Although the field $E \sin \omega t$ appears explicitly in the equation for $h_{\alpha\beta\gamma}$, a transformation similar to those preceding Eqs. (15) and (26) can be made and the field will disappear except for its presence in the new interparticle potential $\phi^{mn} = |\rho_{mn} - \epsilon \sin \omega t|^{-1}$, ϵ being $\epsilon_{\alpha\beta}$, $\epsilon_{\alpha\gamma}$, or $\epsilon_{\beta\gamma}$. For zero field, $|g_{\alpha\beta} / f_{\alpha} f_{\beta}| = 0(N_D^{-1})$; if this ratio remains small for the actual field considered it should be expected that $|h_{\alpha\beta\gamma} / f_{\alpha} g_{\beta\gamma}|$ will remain small too.

We conclude, tentatively, that our results will certainly be valid for $|\underline{\varepsilon}_{ei}|$ small compared with the Debye length, and will probably be valid for $|\underline{\varepsilon}_{ei}|$ comparable to that length if exception is made of the stability restrictions for $\omega \lesssim \omega_p$. (We point out here, finally, that when our results are valid, $\omega_p/\nu(E) \gg 1$ and therefore the expansion made in Sec. II in powers of ν/ω is valid if $\omega/\omega_p \geq 0(1)$.)

We shall discuss, as our last point, the quantitative importance of our result for the function \tilde{Q}_{ii}^0 . The most interesting aspect of the result is the appearance in Eq. (32) of all the resonant factors $(D_{-p})^{-1}$. A detailed qualitative comparison with the moderate field results of Ref. 6 can easily be made. (The moderate field condition is defined by the inequality $v_E \equiv |\omega \underline{\varepsilon}_{ei}| \ll v_e \equiv [kT_e/m_e]^{1/2}$.) In particular, the component of \underline{j}_1 varying as $\sin \omega t$ (the only one heating the plasma) is given in Ref. 6 by

$$\underline{j}_1(\sin \omega t) = \frac{\omega_p^2 Z_e}{\omega 4\pi^3} \int \frac{d\mathbf{k} k}{k^2} \sum_p \frac{J_p^J J_{p+1}^J \text{Im } D_1}{|D_1|^2} \sin \omega t \quad (35)$$

while we include the factor $k^2 [k^2 + k_i^2 \sum_p J_p^2/D_{-p}]^{-1}$ inside the integral.

Retaining up to the third power of E, our integral becomes

$$\int \frac{d\mathbf{k} k}{k^2} \frac{-\underline{\varepsilon}_e \cdot \underline{k}}{2} \left[\frac{\text{Im } D_1}{|D_1|^2} - \frac{(\underline{\varepsilon}_e \cdot \underline{k})^2}{6} \left\{ \frac{2\text{Im } D_1}{|D_1|^2} - \frac{\text{Im } D_2}{|D_2|^2} \right\} \right] \\ \times \left[\frac{k^2}{k^2 + k_i^2 \left\{ \frac{1 - (\underline{\varepsilon}_e \cdot \underline{k})^2/2}{D_0} + \frac{(\underline{\varepsilon}_e \cdot \underline{k})^2}{2} \frac{\text{Re } D_1}{|D_1|^2} \right\}} \right] \quad (36)$$

while the second bracket does not appear in Ref. 6. Here the prefixes Re and Im stand for the real and imaginary parts, and we wrote $\underline{\underline{\varepsilon}}_e$ for $\underline{\underline{\varepsilon}}_{ei}$ since $m_i \rightarrow \infty$. We point out at this stage that the use of the small argument expansion in the Bessel functions in Eq. (35) that we have used to obtain Eq. (36), is valid under the moderate field assumption, even if $|\underline{\underline{\varepsilon}}_e \cdot \underline{\underline{k}}|$ is not small. It is possible to show this by following a procedure used in Refs. 1 and 6: when $|\underline{\underline{k}} \cdot \underline{\underline{\varepsilon}}_e| \geq 0(1)$, we have $\omega/kv_e \ll 1$ from our condition $v_E \ll v_e$; then one solves Eq. (19) by expanding $\tilde{\underline{\underline{Q}}}_{ei}(\tau)$ and $\exp[-ik \cdot \underline{\underline{\varepsilon}}_e \sin \omega\tau]$ inside the time integral in a Taylor series in ω/kv_e and solving iteratively for $\tilde{\underline{\underline{Q}}}_{ei}(t)$. The result for the integral in $J_1(\sin \omega t)$ agrees with Eq. (36) when the small argument expansion is made in $D_p(p\omega/kv_e)$.

A subsequent expansion of the second bracket in Eq. (36) is perhaps not valid throughout the whole range of the $\underline{\underline{k}}$ integration. To make our comparison simpler we shall assume so here. Then this bracket becomes:

$$\frac{k^2 + k_e^2}{k^2 + k_e^2 + k_i^2} \left\{ 1 - \frac{k_i^2}{k^2 + k_e^2 + k_i^2} \left[-\frac{(\underline{\underline{\varepsilon}}_e \cdot \underline{\underline{k}})^2}{2} + 0 \left(\frac{\omega^2 \underline{\underline{\varepsilon}}_e^2}{v_e^2} \right) + \frac{k^2 + k_e^2}{k^2} \right. \right. \\ \left. \left. \times \frac{(\underline{\underline{\varepsilon}}_e \cdot \underline{\underline{k}})^2}{2} \frac{\text{Re}D_1}{|D_1|^2} \right] \right\}$$

The term $0(\omega^2 \underline{\underline{\varepsilon}}_e^2 / v_e^2)$ is due to the fact that $f_{eo} = f_{em} \left[1 + 0(\omega^2 \underline{\underline{\varepsilon}}_e^2 / v_e^2) \right]$; a similar correction with respect to the linear result will appear in the first term inside the first bracket in Eq. (36). Since we retain only third powers of $\underline{\underline{\varepsilon}}_e$, f_{em} can be used in the other terms in Eq. (36) involving D_1 and D_2 .

We see, finally, that to the nonlinear terms given in Ref. 6 [as adding to the linear result inside the integral in Eq. (36)],

$$-\frac{(\underline{\epsilon} \cdot \underline{k})^2}{6} \left\{ \frac{2I_m D_1}{|D_1|^2} - \frac{I_m D_2}{|D_2|^2} \right\} ,$$

we add the essential corrections

$$-\frac{I_m D_1}{|D_1|^2} \frac{(k_e^2 + k_i^2) k_i^2}{(k^2 + k_e^2 + k_i^2)^2} \left[0 \left(\frac{\epsilon_e^2 \omega^2}{v_e^2} \right) - \frac{(\underline{\epsilon} \cdot \underline{k})^2}{2} + \frac{k^2 + k_e^2}{k^2} \frac{(\underline{\epsilon} \cdot \underline{k})^2}{2} \frac{R_e D_1}{|D_1|^2} \right] .$$

In fact, the last term inside this bracket seems to yield the most dominant nonlinear term.

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